

# Trajectory Design and Optimization for LEO Satellites in Formation to Observe GEO Satellites' Beams

## Yi Lu<sup>1</sup>, Yu Sun<sup>1</sup>, Xiyun Hou<sup>2</sup>, Yunhe Meng<sup>3</sup>

<sup>1</sup>State Key Laboratory of Astronautic Dynamics, China Xi'an Satellite Control Center, Xi'an, China <sup>2</sup>School of Astronomy & Space Science, Nanjing University, Nanjing, China <sup>3</sup>Tianqin Research Center for Gravitational Physics, Sun Yat-sen University, Zhuhai, China Email: luyi\_xscc@126.com

How to cite this paper: Lu, Y., Sun, Y., Hou, X.Y. and Meng, Y.H. (2018) Trajectory Design and Optimization for LEO Satellites in Formation to Observe GEO Satellites' Beams. *International Journal of Astronomy and Astrophysics*, **8**, 368-385. https://doi.org/10.4236/ijaa.2018.84026

Received: September 7, 2018 Accepted: December 15, 2018 Published: December 18, 2018

Copyright © 2018 by authors and Scientific Research Publishing Inc. This work is licensed under the Creative Commons Attribution International License (CC BY 4.0).

http://creativecommons.org/licenses/by/4.0/

# Abstract

This paper presents the methods and results for the trajectory design and optimization for the low earth orbit (LEO) satellites in formation to observe the geostationary orbit (GEO) satellites' beams. The background of the trajectory design mission is the 9<sup>th</sup> China Trajectory Optimization Competition (CTOC9). The formation is designed according to the observation demands. The flying sequence is determined by a reference satellite using a proposed improved ephemeris matching method (IEMM). The formation is changed, maintained and transferred following the reference satellite employing a multi-impulse control method (MICM). Then the total observation value is computed by propagating the orbits of the satellites according to the sequence and transfer strategies. Based on the above methods, we have obtained a fourth prize in the CTOC9. The proposed methods are not only fit for this competition, but can also be used to fulfill the trajectory design missions for similar multi-object explorations.

## **Keywords**

Trajectory Design, Global Optimization, Formation Flying, Flying Sequence, Formation Maintaining

# **1. Introduction**

Geostationary orbit (GEO) plays an important role in communication, navigation, and other areas [1] [2]. With an orbital period equal to the Earth's rotation period, the GEO satellites are always above the Earth's equator, and are fixed relative to an Earth-fixed observer [3]. This principal characteristic of the GEO makes it suitable for communication and navigation satellites [4]. Meanwhile, it is essential of the frequency resources for construction of the communication and navigation satellites' systems [5]. Hence, it is of significance to monitor the usage of the frequency resources of GEO satellites. Based on this background, problem A of CTOC9 has considered the orbit design and optimization mission for the low earth orbit (LEO) satellites in formation to observe the GEO satellites' beams [6]. The J2 perturbation must be considered when propagating the orbits and maintaining the flying formation [7].

The problem is a multi-object optimization problem, which appears in many practical applications. For example, the multi-asteroid exploration, the multi-planet exploration and the multi-debris clearing mission [8] [9]. At the end of the CTOC9, teams from NUAA [10], NUDT [11], and Tsinghua University [12] have obtained the first, second, and third prizes. And we have obtained the fourth. Most of them used the regressive orbit to solve the special problem, but our methods are different from theirs. We attempted to propose more general methods to solve not only this problem but also other multi-objective trajectory design problems.

This paper is organized as follows. Section 2 gives a brief description of the problem. Section 3 gives the general description of the methods. Application of the methods to the problem is presented in Section 4. Results and conclusions are finally presented in Sections 5 and 6.

## 2. Brief Description of the Problem

In the problem, the GEO constellation consists of 18 satellites, each of which is equipped with a transmitter which directs its beam to the ground station. The diagram of the projection of the GEO satellites' beams is shown in **Figure 1**.

An LEO formation consisting of 3 monitoring satellites is inserted into circular orbits of altitude of 400 km and inclination of 42.8° at initial time. The task of the three monitoring satellites is to monitor all the beams of the GEO satellites. The "monitoring" operation of a GEO satellite beam is executed only when the three monitoring satellites are in the coverage of the beam. Monitoring profit of a GEO satellite, which is defined by the geometry between the three monitoring execution. Assuming the three monitoring satellites are in the coverage of a GEO satellite's beam at a specified time, **Figure 2** shows the geometry of the three monitoring satellites and the GEO satellite. When the monitoring profits of all 18 GEO satellites reach a certain threshold (In the problem, it is  $1 \times 10^7$  km<sup>2</sup>), the whole monitoring task is completed. Please refer to Ref. [6] for more information of the monitoring profits' computational method and the overall description of the problem.

There are two optimization objectives of different priorities in the problem. The primary objective is to minimize the time duration of the whole monitoring



**Figure 1.** The diagram of the projection of the GEO satellites' beams [6]. (a) Refers to the two-dimensional view and (b) Refers to the three-dimensional view.



Figure 2. Geometry of the three monitoring satellites and one GEO satellite [6].

task. The secondary objective, in case two or more teams achieve the same primary objective, is to minimize the total fuel consumption of the three monitoring satellites.

Each monitoring satellite is subject to the  $J_2$  perturbation. And the motion equation in the J2000 earth centered inertial coordinate is as follows,

$$\begin{cases} \dot{x} = v_{x} \\ \dot{y} = v_{y} \\ \dot{z} = v_{z} \\ \dot{v}_{x} = -\frac{\mu x}{r^{3}} \left( 1 + \frac{3}{2} J_{2} \left( \frac{R_{e}}{r} \right)^{2} \left( 1 - 5 \frac{z^{2}}{r^{2}} \right) \right) \\ \dot{v}_{y} = -\frac{\mu y}{r^{3}} \left( 1 + \frac{3}{2} J_{2} \left( \frac{R_{e}}{r} \right)^{2} \left( 1 - 5 \frac{z^{2}}{r^{2}} \right) \right) \\ \dot{v}_{z} = -\frac{\mu z}{r^{3}} \left( 1 + \frac{3}{2} J_{2} \left( \frac{R_{e}}{r} \right)^{2} \left( 3 - 5 \frac{z^{2}}{r^{2}} \right) \right) \end{cases}$$
(1)

In which,  $\mu$  is the gravitational constant of the earth. *x*, *y*, and *z* refer to the position of the satellites.  $v_x$ ,  $v_y$ ,  $v_z$  refer to the velocity.  $r = \sqrt{x^2 + y^2 + z^2}$  refers to the range from the earth center to the satellite.  $R_e$  is the radius of the earth, and  $J_2$  is the perturbation constant of the earth's oblateness.

The initial masses of all satellites are 500 kg, and the allowed fuel consumption should not be more than 200 kg. The maneuvers of the monitoring satellites are instantaneous changes of the satellites velocities. After each maneuver, the fuel consumption  $\Delta m$  should be evaluated using the Tsiolkovsky equation.

$$\Delta m = m \left( 1 - e^{\frac{-\Delta V}{g_e l_{sp}}} \right)$$
(2)

where *m* is the mass before the maneuver,  $\Delta V$  is the magnitude of the maneuver,  $g_e$  is the acceleration of gravity at sea level of Earth,  $I_{sp}$  is the specific impulse of the propulsion.

The GEO satellites are assumed to follow non-perturbed Keplerian orbits. The longitude and radius of each GEO satellite remain constant, and the latitude remains zero during the whole mission. For more information of the problem description, please refer to Ref. [6].

#### 3. General Description of the Method

#### 3.1. Outline of the Method

In order to solve the problem effectively, we decomposed the methods used by us into following steps.

Step 1: The formation is designed according to the observation demands by analyzing the observation features.

Step 2: The flying sequence is firstly determined by a specified reference satellite using a proposed improved ephemeris matching method (IEMM). Then, the formation is maintained and transferred following the reference satellite by a multi-impulse control method (MICM). The total monitoring profit is computed by propagating the satellites' orbits subject to  $J_2$  perturbation according to the sequence and transfer strategies.

Step 3: The result data is checked and submitted to the competition organizer.

The basic procedure of the methods employed in this paper for solving the problem A of CTOC9 is described in **Figure 3**.

#### 3.2. Formation Designing

In order to design stable formations and reduce the maintaining cost, we analyzed the stability of the formations by propagating the differential equations of the elements difference between the satellites. In the presence of the  $J_2$  perturbation, some orbital elements change, including the drifts in perigee and mean anomaly, and the nodal regression. Employing orbit-averaged quantities [13] [14], differences in the orbital elements between the leader and the follower satellites are given as below:

$$\begin{cases} \delta \dot{a} = \delta \dot{e} = \delta i = 0 \\ \delta \dot{\Omega} = -\frac{3\sqrt{\mu}J_2R_e^2}{2\left(1-e_L^2\right)^2 a_L^{3.5}} \left[ -\sin i_L \delta i + \cos i_L \left(\frac{4e_L \delta e}{1-e_L^2} - \frac{7\delta a}{2a_L}\right) \right] \\ \delta \dot{\omega} = \frac{3\sqrt{\mu}J_2R_e^2}{4\left(1-e_L^2\right)^2 a_L^{3.5}} \left[ -5\sin 2i_L \delta i + \left(5\cos^2 i_L - 1\right) \left(\frac{4e_L \delta e}{1-e_L^2} - \frac{7\delta a}{2a_L}\right) \right] \\ \delta \dot{M} = -\frac{3}{2}\sqrt{\frac{\mu}{a_L^3}} \frac{\delta a}{a_L} + \frac{3\sqrt{\mu}J_2R_e^2}{4\left(1-e_L^2\right)^{1.5} a_L^{3.5}} \\ \times \left[ -3\sin 2i_L \delta i + \left(3\cos^2 i_L - 1\right) \left(\frac{3e_L \delta e}{1-e_L^2} - \frac{7\delta a}{2a_L}\right) \right] \end{cases}$$
(3)

where  $\mu$  is the gravity constant of Earth. *a*, *e*, *i*,  $\Omega$ ,  $\omega$ , *M* are the orbital elements of the satellites, and  $\delta a$ ,  $\delta e$ ,  $\delta i$ ,  $\delta \Omega$ ,  $\delta \omega$ ,  $\delta M$  are the elements' differences between



Figure 3. The procedure of the methods employed for solving the problem A of CTOC9.

the leader and the follower satellites. The elements with subscript L refer to the leader satellites.

Equation (3) shows that the  $J_2$  perturbation does not change the differences of *a*, *e*, *i* but changes the differences of  $\Omega$ ,  $\omega$ , *M* between the leader and the follower satellites. If the initial values of the elements' differences with subscript "0" at time  $t_0$  are given, the values at any time *t* can be computed as follows,

$$\begin{cases} \delta a = \delta a_0, \delta e = \delta e_0, \delta i = \delta i_0 \\ \delta \Omega = \delta \Omega_0 + \delta \dot{\Omega} (t - t_0) \\ \delta \omega = \delta \omega_0 + \delta \dot{\omega} (t - t_0) \\ \delta M = \delta M_0 + \delta \dot{M} (t - t_0) \end{cases}$$
(4)

Equations (3) and (4) obviously show that, when there are differences in the elements a, e or i, the formation is unstable, and changes quickly. As a result, in this paper, the three satellites' initial elements a, e and i are chosen to be nearly the same.

The initial relations of the three satellites' elements will be determined by the mission's requirements, which will be shown in Section 4.1.

#### **3.3. Flying Sequence Searching**

For the flying sequence searching, the improved ephemeris matching method (IEMM) is proposed.

From the characteristics of the Lambert problem [15], it is not hard to understand that, the closer distance from the satellite's position (propagated to the target time without any control from the leaving beam  $b_i$  to the position of the target beam b (propagated to the target time) is, the less energy is needed for the satellite to transfer from  $b_i$  to flyby b. Thereby we can measure the relative size of the transfer energy consumed by examining the closeness between the target beam's ephemeris and the satellite' ephemeris propagated from the leaving beam. In this paper, we define the reference satellite as the leader satellite. When propagating the reference satellite's orbit, the satellite's velocity when leaving the beam is not fixed. We add a small  $\Delta v$  along the velocity in range [-25 m/s, 25 m/s at the departure epoch with a step size of 5m/s and propagate the satellite's trajectory. We pick up the one closest to the matching target beam (if the monitoring profit for this beam is not fulfill the requirement), record the  $\Delta v$  value, the arrival time  $t_{\beta}$  and the position **R** and velocity **V** of the closest point from the central line of the beam to the reference satellite's propagated orbit. This process is different from the ephemeris matching method (EMM) in Ref. [16], in which the trajectory is propagated without  $\Delta v$  along the velocity, so we call it improved ephemeris matching method (IEMM).

For a segment of the transfer trajectories, the mathematical model of IEMM can be expressed as

$$\Delta v = f\left(\lambda_{r} \left\| \boldsymbol{R}(t) - \boldsymbol{R}_{b}(t) \right\|, \lambda_{v} \left\| \boldsymbol{V}(t) - \boldsymbol{V}_{b}(t) \right\| \right)$$
(5)

where  $f(\cdot)$  is a monotonic increasing function; *R* and *V* are respectively the position and velocity of the satellite's propagated orbit at arrival time *t*; *R*<sub>b</sub> and *V*<sub>b</sub> are

respectively the position and velocity of a point on the central line of a beam *b*. The point is chosen to be with the same height to the satellite's propagated orbit at the arrival time.  $\lambda_r$  and  $\lambda_v$  are the weights of the positions' difference and velocities' difference. For this paper, the satellites need to flyby (not rendezvous) the target beam, so the velocities between them need not to be equal. Only the positions need to match. Therefore,  $\lambda_v$  is set to be 0, and  $\lambda_r$  could be 1.

For a transfer segment, when we have set a  $\Delta v_i$  (which is determined by step 5m/s in the small range [-25, 25] m/s from lowest to highest), the target beam  $b_i$  whose ephemeris matches the reference satellite could be found along the propagating satellite's orbit till  $t_i$ . Then we identify the target beam and record the corresponding parameters that meet the condition of top 5 values of the net accumulated monitoring profits when propagating the satellites orbits in the segment. Here net profit means when the accumulated monitoring profit of a GEO satellite's beam has up to the threshold, it will not be added to the total accumulated value computed along this segment. When computing the monitoring profit, the formation is formed theoretically by the geometry relationship between the leader and the follower satellites, because in this sequence searching process, the formation maintaining is not considered.

The procedure of the IEMM algorithm is shown in **Figure 4**, and is described as follows:



**Figure 4.** The procedure of the IEMM algorithm. The solid arrows and circles refer to the recorded sequences, and the dotted arrows and circles refer to the abandoned sequences. The red arrows are connected to show an example sequence *c*.

Step 1: Propagate the reference satellite's orbits from the last beams of the recorded sequences with the initial velocity offset  $\Delta v_i$  from -25 m/s to 25 m/s with a step size of 5 m/s. There will be 11 orbits to be propagated.

Step 2: Find the 11 target beams whose central line is closest to the satellite's orbits corresponding to  $\Delta v_i$  at the end time  $t_i$ .

Step 3: Propagate the reference satellite's orbit and use the theoretical formation (by the geometry relationship between the leader and follower satellites) to compute the net accumulated monitoring profits.

Step 4: For a sequence, if the total monitoring profit meets the condition of all the beams' monitoring profits up to the threshold, record the new sequence as one of the candidate sequences. If any one of the satellite's fuel has all been consumed but not all the observation values of the beams have been up to the threshold, the sequence is abandoned. Otherwise, abandon the 6 sequences with the least monitoring profits, record the top 5 sequences for further searching, and turn to Step 1.

Step 5: If all recorded sequences have fulfilled the threshold, stop searching, and pick out the sequence with shortest mission time for further processing.

#### **3.4. Formation Maintaining**

For the formation maintaining and orbit transfer, a multi-impulse control method (MICM) is employed. The  $\Delta vs$  of all maneuvers are optimized by the differential evolution (DE) algorithm and the  $J_2$  perturbed Lambert problem solving algorithm. The algorithm procedure is shown in **Figure 5**.



Figure 5. The procedure of the MICM algorithm.

There are two layers of the loop. The inner layer is the solving of the  $J_2$  perturbed Lambert problem. The locally optimization shooting method is used to solve the boundary value problem. If the three satellites' final positions  $R_r$  and velocities  $V_r$  equal the values predetermined by the theoretical relationship of the formation, the problem is solved. The predetermined values are given by the sequence searching and formation designing algorithms. When the position and velocity of the reference satellite are provided, based on the relationship between the three satellites in formation, the required positions and velocities of the follower satellites are determined, which are the objective values of the shooting method.

The outer layer of the loop is using the DE algorithm to optimize the velocity increments ( $\Delta vs$ ) to be minimal. The optimizing problem is a time-fixed optimal transfer problem. DE is a heuristic searching algorithm introduced by Storn and Price [17]. Its remarkable performance as a global optimization algorithm on continuous numerical minimization problems has been extensively explored [18]. As with other evolutionary algorithms, DE solves optimization problems by evolving a population of candidate solutions using alteration and selection operators. For the problem in this paper, the objective function is as follows:

$$I = \min\left\{\Delta v_{i,k}\right\} \tag{7}$$

where  $\Delta v_{i,k}$  refers to the sum of the velocity increments of the *t*<sup>h</sup> transfer of satellite *k*. The variables to be determined by the optimization are as follows,

$$\begin{cases} \mathbf{\Phi}_{1,k} = \left[ \Delta t_{1,k}, \Omega_{1,k}, \omega_{1,k}, M_{1,k}, \Delta t'_{1,k}, \Delta R_{1,x}, \Delta R_{1,y}, \Delta R_{1,z} \right] \\ \mathbf{\Phi}_{c,k} = \left[ \Delta t_{c,k}, \Delta t'_{c,k}, \Delta R_{c,x}, \Delta R_{c,y}, \Delta R_{c,z} \right] \\ \mathbf{\Phi}_{i,k} = \left[ \Delta t_{i,k}, \Delta \omega_{i,k} \right], i \neq 1 \neq c \end{cases}$$

$$\tag{8}$$

in which  $\Phi_{1,k}$  indicates the variables relative to the three-impulse transfer for the first maneuver of satellite k;  $\Delta t_{1,k}$  is the waiting time when the satellite is flying along the initial orbit;  $\Omega_{1,k}$ ,  $\omega_{1,k}$ , and  $M_{1,k}$  are respectively the right ascension node, the argument of perigee, and the mean anomaly of the initial orbit;  $\Delta t'_{1,k}$  is the transfer time from the leaving point to a middle point on an initial  $J_2$  perturbed Lambert orbit;  $\Delta R_{1,x}$ ,  $\Delta R_{1,y}$ , and  $\Delta R_{1,z}$  are the position offsets from the middle points. Then two Lambert orbits are computed again to fulfill the transfers from the leaving point to the middle point and from the middle point to the end point.

 $\Phi_{c,k}$  indicates the variables relative to the three-impulse transfer for formation changing (Why the formation need to be changed? Please refer to section 4.1.) of satellite k;  $\Delta t_{c,k}$  is the waiting time, at the end of which the satellite start to transfer;  $\Delta t'_{c,k}$  is the transfer time from the leaving point to a middle point on an initial  $J_2$  perturbed Lambert orbit;  $\Delta R_{c,x}$ ,  $\Delta R_{c,y}$ , and  $\Delta R_{c,z}$  are the position offsets from the middle points. Then two Lambert orbits are computed again to fulfill the transfers from the leaving point to the middle point and from the middle point.

 $\Phi_{i,k}$  indicates the variables relative to the two-impulse transfers for the rest maneuvers of satellite k;  $\Delta \omega_{i,k}$  is the difference of the argument of perigee between the leader satellite and the follower satellite. When propagating the orbits, only small difference of  $\Delta \omega_{i,k}$  will be generated by the  $J_2$  perturbation. In order to save fuel, we do not require the argument of perigee for the 3 satellites exactly equal each other. But when solving the  $J_2$  perturbed Lambert problem, the offset of the target mean anomaly should be set to have an opposite value to  $\Delta \omega_{i,k}$ . This will ensure a relatively stable shape to fulfill the observation. The bounds of the variables are shown in **Table 1**.

#### 3.5. Monitoring Profit Computation

Incorporating all the maneuvers of the 3 monitoring satellites, the observation values  $W_j$  are computed by propagating the  $J_2$  perturbed orbits with a step time  $t_m$  of 1 second using Equation (9) [6]:

$$W_{j} = \sum_{m=1}^{M} S_{j}\left(t_{m}\right) g_{j}\left(t_{m}\right)$$
(9)

where *j* refers to the label of the GEO satellite's beam,  $S_j$  is the area of the triangle formed by 3 monitoring satellites, and  $g_k$  is sine of the angle from the GEO satellite's beam to the triangle plane. The results are organized to fit the submission format in order to pass the data checking process. *M* refers to the total duration (in second) when the triangle is in the beam *j*.

## 4. Application of the Method to the Problem

## 4.1. Formation Determination

In order to monitor the 11<sup>th</sup> GEO satellite whose beam is locating at the highest latitude (which is shown in **Figure 6**), we raise the apogees of the three orbits and make the formation be a triangle parallel with the orbit plane to reach the beam. Of course if we raise the inclinations of the orbits, the 11<sup>th</sup> beam could also be monitored, but it will consume too much fuel. So we choose the first way.

	$\Delta t_{1,k}/s$	$\Omega_{1,k}/\text{deg}$	$\omega_{1,k}/\deg$	M <sub>1,k</sub> /deg	$\Delta t'_{1,k}/s$	$\Delta R_{1,x}/\mathrm{km}$	$\Delta R_{1,y}/\mathrm{km}$	$\Delta R_{1,z}/\mathrm{km}$
up	$\Delta T_1/2$	10	90	90	90	30	30	30
low	0	-10	-90	-90	-90	-30	-30	-30
	$\Delta t_{c,k}/s$	$\Delta \omega_{c,k}/\deg$	$\Delta t'_{\rm c,k}/s$	$\Delta R_{c,x}/km$	$\Delta R_{c,y}/\text{km}$	$\Delta R_{\rm c,z}/{\rm km}$		
up	180	4	$\Delta T_c + 500$	50	50	50		
low	-60	-4	$\Delta T_c-500$	-50	-50	-50		
	$\Delta t_{i,k}/s$	$\Delta \omega_{i,k}/\text{deg}$						
up	45	2						
low	-45	-2						

Table 1. Bounds of the variables.

Note:  $\Delta T_1$  is the time range from initial to the first beam in the sequence.  $\Delta T_c$  is the time range from the leaving to the arriving point of the formation changing transfer.



**Figure 6.** The three-diminsional geometry of the 11<sup>th</sup> beam, which is the highest one in the figure. The yellow thin lines are the satellites's initial orbits.

When the monitoring profit of the  $11^{\text{th}}$  GEO satellite's beam is near  $1 \times 10^7$  km<sup>2</sup>, the formation is changed from the first one to the second one which constitutes a triangle perpendicular to the orbit plane for the observing of the remaining GEO satellites' beams locating at the low latitudes. The first and the second formations are shown in **Figure 7**.

When designing the formations, not only the monitoring requirements, but also the stability should be considered to reduce the formation maintaining cost. Finally, the relations of the 3 satellites are determined and shown in **Table 2**.

The leader satellites' elements  $a_D e_D i_D \Omega_D \omega_D M_L$  are determined by the flying sequence and transferring strategies. The follower satellites are transferred to fit the relations in **Table 2**. The values of differences between the elements of the leader and the followers are determined by optimization to find the highest mean monitoring profit for passes of the GEO satellite's beams.

## 4.2. Flying Sequence Determination

For the flying sequence determination, the proposed IEMM described in Section 3.3 is employed. For every searched sequence, the labels of the beams, the epochs when the leader satellite arrive the beams and the positions & velocities of the orbits are recorded. Then the top 5 beams with high monitoring profits are chosen and added to the sequence to form new sequences. For a special sequence, if the total monitoring profit meets the condition of all the beams' monitoring profits up to the threshold, record the new sequence as one of the candidate sequences. If any one of the satellite's fuel has all been consumed but not all the monitoring profits of the beams have been up to the thresholds, the sequence is abandoned. Otherwise, abandon the last 6 values, and record the top 5 sequences for further searching. If all recorded sequences have fulfilled the threshold condition, stop searching, and pick out the sequence with shortest mission time for further processing.

The sequence searching result is shown in **Figure 8**, which shows that most of the durations between two target beams of the ephemeris matching processes are no longer than  $1 \times 10^5$  s (27.8 hours). Most of the distances between the



**Figure 7.** The designed formations. The red star refers to the reference satellite which is also the leader satellite, and the other two are the follower satellites. (a) The first formation constitutes a triangle parallel to the orbit plane. (b) The second formation constitutes a triangle perpendicular to the orbit plane.



**Figure 8.** The sequence searching result. The first sub-figure refers to the beam numbers of the best sequence. The second sub-figure refers to the duration between two points of the sequence. The third sub-figure refers to the distance between the reference satellite's orbit and the central line of the corresponding target beam. And the fourth sub-figure refers to the change of the reference satellite's mass.

<b>Table 2.</b> The relations of the 3 satellites
---

Formation	Satellite	а	е	i	Ω	ω	M
	Leader	$a_L$	$e_L$	$i_L$	$arOmega_L$	$\omega_L$	$M_L$
Formation 1	Follower 1	$a_L$	$e_L$	$i_L$	$arOmega_L$	$\omega_L$	$M_L - 2.237$
	Follower 2	$a_L$	$e_{L} - 0.02$	$i_L$	$\Omega_L + 3$	$\omega_L$	$M_L$
	Leader	$a_L$	$e_L$	$i_L$	$arOmega_L$	$\omega_L$	$M_{\scriptscriptstyle L}$
Formation 2	Follower 1	$a_L$	$e_L$	$i_L$	$arOmega_L$	$\omega_L$	$M_L$ + 2.237
	Follower 2	$a_L$	$e_L$	$i_L$	$\Omega_L + 3$	$\omega_L$	$M_{L} - 1$

reference satellite's orbits and the central line of the corresponding target beams are no more than 100 km, which means that most of the target beams in addition to the passed beams along the transfer orbit could all be observed by the formation. The consumed fuel by the reference satellite is lower than 160 kg, which means 40 kg left could be used to guarantee the fulfillment of the formation transfer and maintenance along the overall task.

## 4.3. Formation Transfer and Maintenance

For the formation transfer and maintenance, the MICM described in section 3.4 is employed. The time points,  $\Delta vs$  of all maneuvers are optimized by the DE algorithm and  $J_2$  perturbed Lambert problem solving algorithm. Therefore, the satellites can fly following the reference satellite and the formation can be maintained. The orbit of the reference satellite is shown in **Figure 9**, in which the red pluses refer to the maneuvers.

The changes of the reference satellite's elements are shown in **Figure 10**, in which the semi-major and the orbital eccentricity are firstly raised by the initial transfer, and then oscillate in the following transfers. The orbital inclination nearly unchanged to save fuel. And the right ascension node changes from about 225 deg to 100 deg due to the  $J_2$  perturbation.

The consumed fuel for all the maneuvers are shown in **Figure 11**. For every transfer, the velocity increment and the consumed fuel are minimized by the DE algorithm. There are 3 transfers whose consumed fuels are much more than others. The first one is for the transfer from the initial circular orbit to the elliptical orbit, the second one is for the formation changing, and the third one is for the final flying by to monitor the last beam in order to make the total time to be the least while consuming nearly all the fuel of each satellite.



Figure 9. The orbit of the reference satellite (The red pluses refer to the maneuvers).



Figure 10. The changing of the reference satellite's elements.



Figure 11. The consumed fuels for the satellites' maneuvers.

#### 4.4. Monitoring Profit Computation

Incorporating all the maneuvers of the 3 satellites, the monitoring profits (observation values) shown in **Figure 12** are computed using Equation (9) by propagating the orbits. **Figure 12** also shows that all the monitoring profits of the GEO satellites' beams are larger than the threshold of  $1 \times 10^7$  km, which fulfills the requirement of the overall monitoring task.



**Figure 12.** The monitoring profits of all the GEO satellite's beams (The red line refers to the value of  $1 \times 10^7$  km<sup>2</sup>). The distance between every two pluses for every beam refers to the incremented monitoring profit for every monitoring mission.

## **5.** Competition Results

We have obtained the fourth prize of CTOC9. The total flying time is 26.9611 days, and the total consumed fuel is 549.3799 kg, as shown in **Table 3**. The reasons for the discrepancies of the results obtained in this paper and that obtained by the previous research teams at NUAA, NUDT and Tsinghua University may be the amount of calculation is much larger than them, thus missed the optimal solution in a limited duration of the competition.

The initial parameters, the number of maneuvers, and the final consumed fuels of the 3 satellites are shown in **Table 4**.

The final observation values are shown in Table 5.

## 6. Conclusion

This paper presented the methods and results for the problem A of CTOC9, the trajectory design and optimization problem for the low earth orbit (LEO) satellites in formation to observe the geostationary orbit (GEO) satellites' beams. Two methods have been proposed to respectively solve the sequence searching and formation transfer & maintenance problems. The first method is the IEMM and the second one is MICM. Compared with other groups with better results, the methods in this paper may be more powerful in solving similar or more complex multi-objective trajectory design problems, because most of the methods used by other groups employed the regressive orbit to solve the special problem A of CTOC9. If the beams pointing to the surface of the earth to be

Rank	Team Name	1st Object/day	2nd Object/kg
1	Nanjing University of Aeronautics and Astronautics (NUAA)	20.42817	528.15235
2	Research Center of Flight Dynamics and Control (RCFDC), NUDT	21.40815	450.64002
3	Tsinghua University (TU)	24.97238	588.02681
4	Xi'an Satellite Control Center (XSCC)	26.96105	549.37992
5	Northwestern Polytechnical University (NPU)	27.22356	566.65516
6	Laboratory of Space System Dynamics and Simulation Technology (LSSDST), NUDT	29.06861	545.47487
7	China Academy of Launch Vehicle Technology(CALVT)	41.53081	423.36019
8	China Academy of Space Technology (CAST)	48.96269	526.07781

**Table 3.** The results of the objective values. (To make a clear comparison, we also presented the results of the other teams in the table. The contents in bold is our result.)

 Table 4. Initial parameters and the final consumed fuels.

	Leader	Follower 1	Follower 2
<i>a</i> (km)	6778.0676	6778.1642	6778.1172
е	$1.1058\times10^{-5}$	$4.0258\times10^{-6}$	$5.5090 \times 10^{-6}$
i(deg)	42.8000	42.8000	42.8000
$\Omega$ (deg)	224.6348	224.5748	227.6658
$\omega$ (deg)	67.5210	274.2653	18.3462
$M(\deg)$	202.4822	358.0967	237.9293
Number of maneuvers	116	116	116
Consumed fuel	175.8506	177.3275	196.2017

#### Table 5. The final monitoring profits.

GEO beam	Monitoring profit/1 $\times$ 10 $^{7}km^{2}$	GEO beam	Monitoring profit/1 $\times$ $10^7km^2$
1	1.3684	10	1.0117
2	1.2856	11	1.1730
3	1.0462	12	1.2073
4	1.2803	13	1.4633
5	1.1974	14	1.0422
6	1.0284	15	1.0859
7	1.2502	16	1.2151
8	1.2091	17	1.2087
9	1.0815	18	1.2898

monitored are not fixed, the regressive orbit may not be an effective candidate tool to solve the problem. But we have attempted to propose general methods

which need not any special characteristics of the monitor or the target. The proposed methods may have a wider application area to solve more multi-objective trajectory design and optimization problems.

## Acknowledgements

We are very grateful to the 9<sup>th</sup> China trajectory optimization competition organizers for the interesting problem, challenge, and fruitful discussions. This work was supported by the Youth Fund Project of the Space Systems Department (No. 2017SY26B0001), the Funds of Natural Science of China (No. 61673390, and No. 61503414), and the Fund of The State Key Laboratory of Astronautic Dynamics (No. 2016ADL-DW0202).

## **Conflicts of Interest**

The authors declare no conflicts of interest regarding the publication of this paper.

#### References

- [1] Shi, R., Liu, L., Long, T., *et al.* (2017) Surrogate Assisted Multidisciplinary Design Optimization for an All-Electric GEO Satellite. *Acta Astronautica*, **138**, 301-317.
- [2] Asgarimehr, M. and Hossainali, M.M. (2014) Optimization of Geosynchronous Satellite Constellation for Independent Regional Navigation and Positioning in Middle East Region. *Acta Astronautica*, 104, 147-158. https://doi.org/10.1016/j.actaastro.2014.07.025
- Pocha, J.J. (1987) An Introduction to Mission Design for Geostationary Satellites. Vol. 6, D. Reidel Publishing Company, Dordrecht, 1-8. https://doi.org/10.1007/978-94-009-3857-1
- [4] Clarke, A.C. (1945) Extra-Terrestrial Relays—Can Rocket Stations Give World-Wide Radio Coverage? *Wireless World*, 305-308.
- [5] Li, H.N. (2014) Geostationary Satellites Collocation. Springer, Berlin Heidelberg. https://doi.org/10.1007/978-3-642-40799-4
- [6] Zhao, S.G., Qi, R., Zhang, J.R., *et al.* (2018) Problem A of 9th Trajectory Optimization Competition: Problem Description and Summary of the Results. *Acta Astronautica.*
- [7] Alfriend, K.T. and Schaub, H. (2000) Dynamics and Control of Spacecraft Formations: Challenges and Some Solutions. *Journal of the Astronautical Sciences*, 48, 249-267.
- [8] Zhu, K., Jiang, F., Li, J., et al. (2009) Trajectory Optimization of Multi-Asteroids Exploration with Low Thrust. Transactions of the Japan Society for Aeronautical and Space Sciences, 175, 47-54. <u>https://doi.org/10.2322/tjsass.52.47</u>
- Bérend, N. and Olive, X. (2016) Bi-Objective Optimization of a Multiple-Target Active Debris Removal Mission. *Acta Astronautica*, **122**, 324-335. https://doi.org/10.1016/j.actaastro.2016.02.005
- [10] Yang, B., Huang, X.X., Yang, H.W., et al. (2018) Problem A of 9th China Trajectory Optimization Competition: Results Found at NUAA. Acta Astronautica.
- [11] Wu, D., Song, Y., Chi, Z.M., *et al.* (2018) Problem A of the 9th China Trajectory Optimization Competition: Results Found at Tsinghua University. *Acta Astronau*-

tica.

- [12] Liang, J., Yang, Z., Zhang, J., *et al.* (2018) Formation Establishment for GEO Satellite Beams Monitoring Using Controlled Regressive Orbits.
- [13] Sabol, C., Burn, R. and McLaughlin, C.A. (2001) Satellite Formation Flying Design and Evolution. *Journal of Satellite and Rockets*, 38, 270-278. https://doi.org/10.2514/2.3681
- [14] Li, J.F., Meng, X., Gao, Y.F., *et al.* (2005) Study on Relative Orbital Configuration in Satellite Formation Flying. *Acta Mechanica Sinica*, 21, 87-94. <u>https://doi.org/10.1007/s10409-004-0009-3</u>
- [15] Curtis, H.D. (2005) Orbital Mechanics for Engineering Students. Elsevier, 163-172.
- [16] Lu, Y., Li, J., Li, H.N., et al. (2014) Fast Design of Low-Thrust Transfer Orbit for Manned Asteroids Exploration. 65th International Astronautical Congress, Toronto, 2-6 October 2014.
- [17] Storn, R. and Price, K. (1997) Differential Evolution—A Simple and Efficient Heuristic for Global Optimization over Continuous Spaces. *Journal of Global Optimization*, 11, 341-359. https://doi.org/10.1023/A:1008202821328
- [18] Price, K.V., Storn, R.M. and Lampinen, J.A. (2006) Differential Evolution: A Practical Approach to Global Optimization. 2nd Edition, Springer-Verlag, Berlin, 37-134.