

# Modification of Even-A Nuclear Mass Formula

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## Abstract

In this paper we obtain an empirical mass formula of even-A nuclei based on residual proton-neutron interactions. The root-mean-squared deviation (RMSD) from experimental data is at an accuracy of about 150 Kev. While for heavy nuclei, we give another formula that fits the experimental data better (RMSD  $\approx$  119 Kev). We have successfully described the experimental data of nuclear masses and predicted some unknown masses (like  $^{200}\text{Ir}$  not involved in AME2003, the deviation of our predicted masses from the value in AME2012 is only about 82 keV). The predictive power of our formula is more competitive than other mass models.

## Keywords

Residual Proton-Neutron Interactions, Nuclear Masses, Binding Energies

## 1. Introduction

The study of nuclear masses and energy levels has always been one of the most challenging frontiers in the field of nuclear physics. There are two types to describe and understand the nuclear masses, one of which is global relations, and the other is local. Some global nuclear mass models such as Weizsäcker model [1], Duflo-Zuker model [2], the finite range droplet model [3], a recent macroscopic-microscopic mass formula [4] [5] [6] etc., successfully produce the measured masses with accuracy at the level of 300 - 600 Kev. However, the global mass models require more physics and more information about nuclear force to get better description of the nuclear masses. On the other hand, the local mass relations, such as the isobaric multiplet mass equation (IMME), the Garvey-Kelson (GK) relations, which use the predicted nuclear masses and the residual proton-neutron interactions to evaluate the mass. It is found that the local mass relations are just approximately satisfied in known masses, so it has a good potential to predict the unknown masses.

In this paper, our purpose is to obtain a residual proton-neutron interactions formula of even- $A$  nuclei from those of neighboring nuclei. In Section II we introduce the residual proton-neutron interactions and obtain our formula based on the proton-neutron interactions between the last proton and the last neutron. Then we introduce two modifications to improve our formula. The RMSD from experimental data is about 150 Kev. And for heavy nuclei, we obtain another formula fits with the experimental data even more precise. With our further refinement of heavy nuclei, the RMSD gets even smaller to about 120 Kev. In Section III we successfully predict some unknown masses. The result shows that the predict power of our formula is competitive with others. In Section IV we discuss and summarize the results of this paper.

## 2. The Residual Proton-Neutron Interactions

The residual proton-neutron interaction plays an important role in the evolution of collective, deformation and phase transition [7] [8] [9] [10], so it has attracted many attentions [11]-[17]. The proton-neutron interactions between the last  $i$  protons and  $j$  neutrons is given by

$$V_{ip-jn}(Z, N) = B(Z, N) + B(Z - i, N - j) - B(Z, N - j) - B(Z - i, N). \quad (1)$$

The famous formula GKL and GKT were derived from the neutron-proton interactions between the last neutron and proton [18] [19]. The relationship between Garvey-Kelson quality is a semi empirical relationship between 6 adjacent nuclear mass. If the interaction between neighboring nuclei changes slowly in the local range, it can be completely counteracted by the addition and subtraction of many adjacent nuclei. Garvey-Kelson mass relationship has two common relationships:

$$\begin{aligned} M(N, Z + 1) + M(N - 1, Z - 1) + M(N + 1, Z) \\ - M(N, Z - 1) - M(N - 1, Z) - M(N + 1, Z + 1) = 0, \end{aligned} \quad (2)$$

$$\begin{aligned} M(N, Z - 1) + M(N - 1, Z + 1) + M(N + 1, Z) \\ - M(N, Z + 1) - M(N - 1, Z) - M(N + 1, Z - 1) = 0, \end{aligned} \quad (3)$$

where  $M(N, Z)$  denotes the mass of a nucleus with neutron number  $N$  and proton number  $Z$ . Equation (2) is called the longitudinal Garvey-Kelson relation (GKL), and Equation (3) the transverse (GKT).

In this section, we use the residual proton neutron interactions between the last proton and the last neutron to form our formula. According to the Equation (1), it is easy to obtain that the residual proton-neutron interactions between the last proton and the last neutron is defined as

$$\begin{aligned} V_{1p-1n}(Z, N) &= B(Z, N) + B(Z - 1, N - 1) - B(Z, N - 1) - B(Z - 1, N) \\ &= M(Z, N) + M(Z - 1, N - 1) - M(Z, N - 1) - M(Z - 1, N) \end{aligned} \quad (4)$$

The Garvey-Kelson mass relations require six nuclei, but our formula requires only four. So our formula involves less number of nuclei, its predictions in iterative extrapolations is the more reliable, and its deviations are smaller in the

extrapolation process.

In recent years, many papers tried to find formulas to describe and evaluate the nuclear masses, but many of them have a large RMSD. In this work, we focus on the even- $A$  nuclei, through the study on the neighboring nuclei with the database in AME2012 [20].

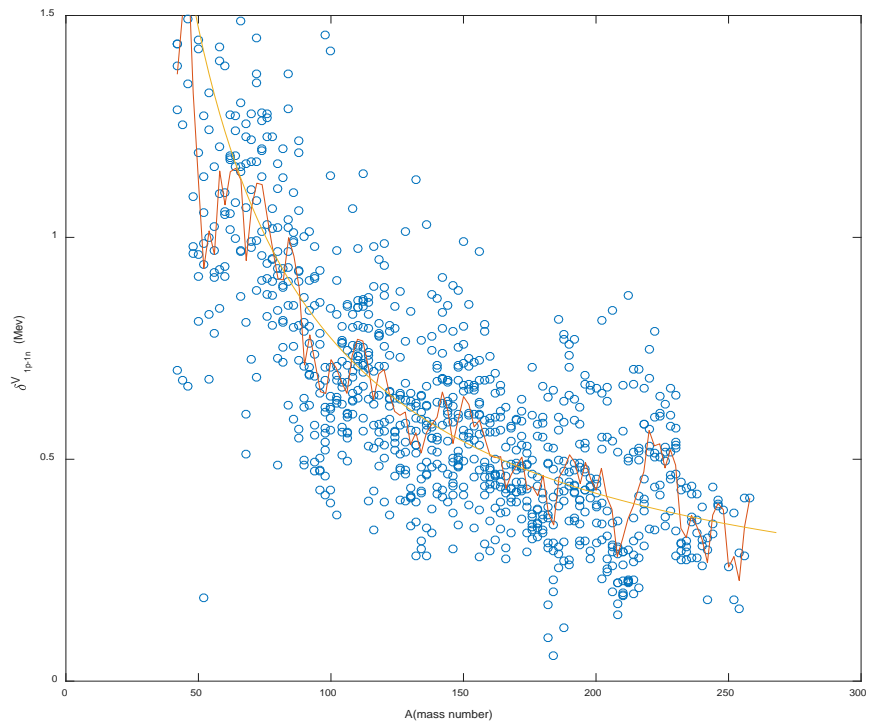
For the residual nuclear proton-neutron interactions which  $A \geq 42$ , we calculate the  $\delta V_{1p-1n}$  as shown in **Figure 1**. Based on that, we empirically obtained the residual proton-neutron interactions formula of even- $A$  nuclei. The formula is as follows:

$$\begin{aligned}\overline{\delta V_{1p-1n}} &= B(Z, N+1) + B(Z-1, N) - B(Z, N) - B(Z-1, N+1) \\ &\cong \frac{515.6}{A^2} + \frac{62.78}{A} + 0.1079 \text{ keV}\end{aligned}\quad (5)$$

$\overline{\delta V_{1p-1n}}$  is the average values of  $\delta V_{1p-1n}$  for nuclei with the same mass number  $A$ .

We find that the average binding energy of our predicted mass agrees well with the specific binding energy curve. We successfully describe and predict some even- $A$  nuclear masses by using these equations and some known experimental nuclear masses in AME2012 for calculation of  $\delta V_{1p-1n}$ .

It can be seen from the **Figure 1** that the interaction of proton-neutron is



**Figure 1.** Circles show that the residual proton-neutron interactions  $\delta V_{1p-1n}$ . The curve is plotted by using the average values of  $\delta V_{1p-1n}$  for nuclei with the same mass number  $A$ , expressed as  $\overline{\delta V_{1p-1n}}$ . The smoothed curve are plotted in terms of equation

$$\overline{\delta V_{1p-1n}}(A) = \frac{515.6}{A^2} + \frac{62.78}{A} + 0.1079 \text{ keV} \quad \text{for even-}A \text{ nuclei with } A \geq 42.$$

more stable in the heavy nuclei region than in the light nuclei region.

In order to better describe the quality of the nucleus, we will improve the above formula with some amendments, donated by  $\delta V_{1p-1n}^{cal}$  as the final improvement results [4] [5] [6]. The first is called the Coulomb correction, denoted by  $\Delta_C$ :

$$\Delta_C(Z, N) \approx a_C \left( -\frac{4}{9} Z^{4/3} A^{-7/3} - \frac{2}{3} Z A^{-4/3} + \frac{4}{9} Z^2 A^{-7/3} + \frac{4}{9} Z^{1/3} A^{-4/3} \right),$$

the second is called the symmetry energy correction, denoted by  $\Delta_{sym}$ :

$$\Delta_{sym}(Z, N) = a_{sym} \frac{1}{A(2 + |IA|)^3} + b_{sym} A^{-1},$$

where  $I = (N - Z)/A$  and  $a_C = 10.51$ ,  $a_{sym} = 20126$ ,  $b_{sym} = -61.25$  as parameters [17] [21].

The revised  $\delta V_{1p-1n}(Z, N)$  is as follows:

$$\delta V_{1p-1n}^{cal}(Z, N) = \overline{\delta V_{1p-1n}} - \Delta_C(Z, N) - \Delta_{sym}(Z, N). \quad (6)$$

The improvement of these two corrections on our predicted  $\delta V_{1p-1n}$  is about 5 keV. Although the two contributions are small, but with more understanding of the symmetry energy of the nucleus, we believe that these contributes will become more important in the future.

In order to describe the nuclear mass obtained by our theory vividly, we compare the average RMSD of the nuclear mass with the experimental data to represent the difference, and the formula is as follows:

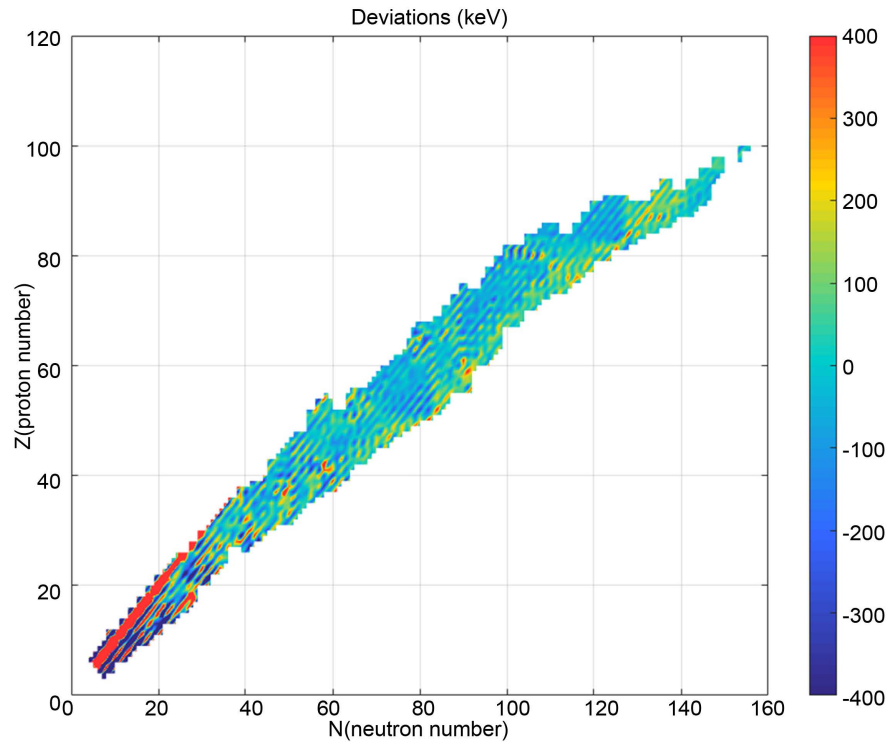
$$\sigma = \sqrt{\frac{1}{n} \sum_{i=1}^n (M_i^{exp} - M_i^{cal})^2}.$$

The RMSD is about 150 Kev. In **Figure 2** we show deviations (in units of keV) between our calculated  $\delta V_{1p-1n}^{cal}$  by applying Equations (6) and those experimental data of binding energies compiled in AME2012 [20]. It can be seen that the RMSDs of these  $\delta V_{1p-1n}$  decrease with  $A$ . The description is better in the medium mass nucleus and heavy nucleus.

As early as 1960s, the nuclear structure theory predicts the existence of a number of new elements in the long life near the proton number  $Z = 114$  and neutron number  $N = 184$  (*i.e.* island of super heavy nuclei) and the island of super heavy nuclear plays an important role in the entire nuclear physics field. So for the heavy nuclei, we obtain another formula to describe the mass and it fits more closely with the experimental data. And in order to achieve better result, the different parameters are given between even-even nuclei and odd-odd nuclei, the formula is as follows:

$$V_{1p-1n}(A) = \frac{a}{A^2} + \frac{b}{A} + c. \quad (7)$$

Parameter	$a$	$b$	$c$
Even-even	-9464	146.3	-0.06435
Odd-odd	46000	-324.1	0.9124



**Figure 2.** (Color online) Deviations (in units of keV) of our calculated  $\delta V_{1p-1n}^{cal}$  by using Equations (6) with respect to those extracted from experimental binding energies [Equation (4)], for the nuclei with  $A \geq 16$ .

When we use the Equation (6) to describe the nuclear masses, the RMSD is about 150 Kev, but if we try the Equation (7) where  $A > 200$ , the RMSD is 119 Kev, it shows that our formula of heavy nuclei is more accurate.

**Figure 3** displays the difference between the experimental values and calculated values, we compare it with Ref [21], one can see that our result is better.

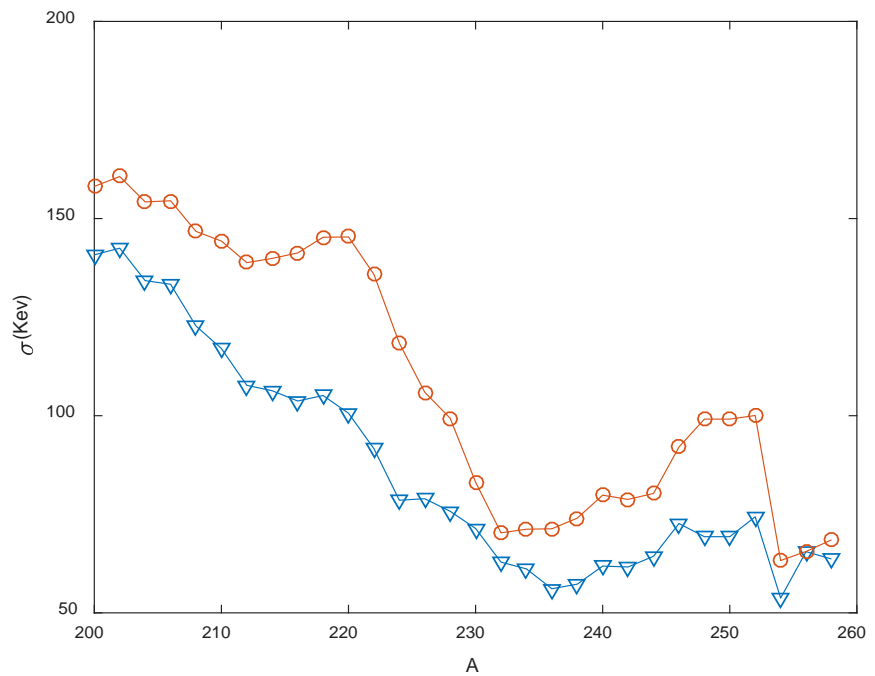
### 3. Mass Predictions

Through above study, we find our formula has a good performance in describing the nuclear masses. In this section, we use our formula and the residual proton-neutron interaction to predict the nuclear mass not obtained in the experiment. Based on the Equation (4), we can obtain

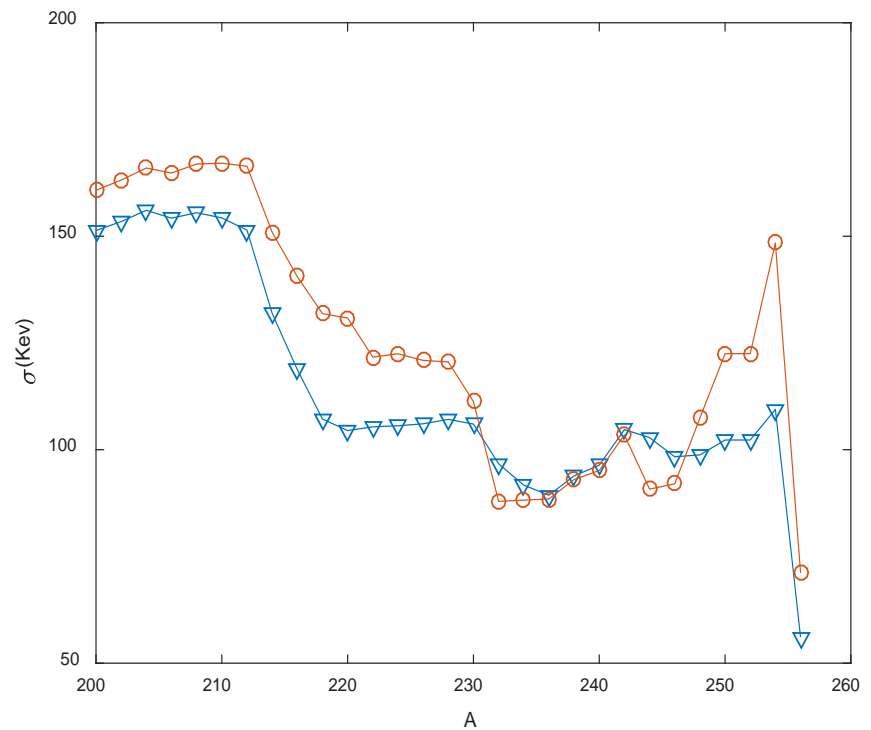
$$M(Z, N) = M(Z-1, N) + M(Z, N-1) - M(Z-1, N-1) + \overline{\delta V_{1p-1n}}(A).$$

The unknown mass  $M(Z, N)$  is predicted by using the three nuclei masses around it and the  $\delta V_{1p-1n}(Z, N)$  we empirical obtained.

Now let's focus on a few examples of our predictions. **Table 1** shows mass excess of some nuclei are not predictive in ame 2003 or ame 2012 databases. These unknown masses are important not only in the context of astrophysics, but also in the nuclear structure. Interestingly, our predicted values show good in comparison with the experimental results. For  $^{182}\text{Lu}$ , the deviation of our predicted masses from the value in AME2012 is only  $\sim 63$  keV. Three additional



(a)



(b)

**Figure 3.** Shows the RMSDs of even-A nuclei. (a) represents the odd-odd nuclei; (b) represents the even-even nuclei. We obtain the even-A nuclear masses from some experimentally known nuclear masses and the residual proton-neutron interactions formula. Comparing calculated values with the AME2012 databases obtain the RMSDs. The triangles are plotted by using the RMSDs of our calculated values. The circles are plotted by using the formula in Ref [21].

**Table 1.** Mass excess of some mass nuclei with us and predicted results in the AME2003 database and the AME2012 databsae. (keV).

Nucleus	AME2003	AME2012	$M^{pred}$
$^{52}\text{Ni}$	-22,650	-23,470	-23,187
$^{74}\text{Sr}$	-40,700	-40,830	-40,952
$^{86}\text{As}$	-59,150	-58,962	-58,316
$^{98}\text{Kr}$	-44,800	-44,310	-44,555
$^{126}\text{Pr}$	-60,260	-60,320	-60,573
$^{148}\text{Tm}$	-39,270	-38,765	-38,713
$^{164}\text{Re}$	-27,640	-27,523	-27,422
$^{182}\text{Lu}$	-41,880	-41,880	-41,817
$^{190}\text{At}$	null	null	10,290
$^{200}\text{Ir}$	null	-21,611	-21,693
$^{202}\text{Pt}$	-22,600	-22,692	-22,592
$^{224}\text{Np}$	null	31,876	31,793
$^{232}\text{Am}$	43,400	43,268	43,376
$^{272}\text{Mt}$	133,890	133,582	133,671
$^{286}\text{Ed}$	168,120	169,725	169,700

nuclei are  $^{202}\text{Pt}$ ,  $^{232}\text{Am}$  and  $^{286}\text{Ed}$ , the differences between our predicted values and those in AME2012 are approximately 100 keV. It seems our formula shows a great accuracy and can be used predict nuclear masses.

#### 4. Discussion and Conclusions

In this paper, we obtain the residual proton-neutron interactions formula to describe and predict the mass of even- $A$  nuclei. In order to improve the accuracy of the  $\delta V_{1p-1n}$ , we use the average value of the  $\delta V_{1p-1n}$  (denoted as  $\overline{\delta V}_{1p-1n}$  modification) and introduce two modifications.

For further understanding of the super heavy nuclei, we use another formula to describe the  $\delta V_{1p-1n}$ , and its results fit the experiment data more accurate, one can see that the RMSD decreases considerably.

Then we investigate the predictive power of these new formulas by numerical experiments. They are competitive with other local mass relations. The deviation of predicted results from experimental values is less compared with other models.

Based on results so far, our method of studying the neighboring nuclei has a good performance. We can predict other unknown masses by using our empirical formula to provide useful reference points for experimental physics.

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