

Penetrative Bénard-Marangoni Convection in a Micropolar Ferrofluid Layer via Internal Heating and Submitted to a Robin Thermal Boundary Conditions

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Abstract

Penetrative Bénard-Maranagoni convection in micropolar ferromagnetic fluid layer in the presence of a uniform vertical magnetic field has been investigated via internal heating model. The lower boundary is considered to be rigid at constant temperature, while the upper boundary free open to the atmosphere is flat and subject to a convective surface boundary condition. The resulting eigenvalue problem is solved numerically by Galerkin method. The stability of the system is found to be dependent on the dimensionless internal heat source strength N_s , magnetic parameter M_1 , the non-linearity of magnetization parameter M_3 , coupling parameter N_1 , spin diffusion parameter N_3 and micropolar heat conduction parameter N_5 . The results show that the onset of ferroconvection is delayed with an increase in N_1 and N_5 but hastens the onset of ferroconvection with an increase in M_1 , M_3 , N_3 and N_{s} . The dimension of ferroconvection cells increases when there is an increase in M_3 , N_1 , N_5 and N_s and decrease in M_1 and N_3 .

Keywords

Bénard-Maranagoni, Micropolar Ferrofluid, Galerkin Method, Penetrative Convection, Internal Heating

1. Introduction

Ferrofluids are colloidal suspensions of magnetic nanoparticles, as suggested by Rosensweig [1] in his monograph, it is pertinent to consider the effect of micro-rotation of the particles in the study. Based on this fact, studies have been undertaken by treating ferrofluids as micropolar fluids and the theory of micropolar fluid proposed by Eringen [2] has been used in investigating the problems. Micropolar fluids have been receiving a great deal of interest and research focus due to their applications like solidification of liquid crystals, the extrusion of polymer fluids, cooling of a metallic plate in a bath colloidal suspension solutions and exotic lubricants. In the uniform magnetic field, the magnetization characteristic depends on particle spin but does not on fluid velocity: Hence micropolar ferrofluid stability studies have become an important field of research these days. Although convective instability problems in a micropolar fluid layer subject to various effects have been studied extensively, the works pertaining to micropolar ferrofluids are in much-to-be desired state. Many researchers (Lebon and Perez [3], Payne and Straughan [4], Siddheshwar and Pranesh [5], Idris et al. [6], Mahmud et al. [7], Sharma and Kumar [8]) have been rigorously investigated the Rayleigh-Bénard situation in Eringen's micropolar non-magnetic fluids. From all these studies, they mainly found that stationary convection is the preferred mode for heating from below. Zahn and Greer [9] have considered interesting possibilities in a planar micropolar ferromagnetic fluid flow with an AC magnetic field. Abraham [10] has investigated the problem of Rayleigh-Bénard convection in a micropolar ferromagnetic fluid layer permeated by a uniform magnetic field for stress-free boundaries. Thermal instability problem in a rotating micropolar ferrofluid has also been considered by Sunil et al. [11]. Nanjundappa *et al.* [12] have investigated the onset of ferromagnetic convection in a micropolar ferromagnetic fluid layer heated from below in the presence of a uniform applied vertical magnetic field.

The practical problems cited above require a mechanism to control thermomagnetic convection. One of the mechanisms to control (suppress or augment) convection is by maintaining a non-uniform temperature gradient across the layer of ferrofluid. Such a temperature gradient may arise due to 1) uniform distribution of heat sources 2) transient heating or cooling at a boundary, 3) temperature modulation at the boundaries and so on. Works have been carried out in this direction but it is still in much-to-be desired state. Rudraiah and Sekhar [13] have investigated convection in a ferrofluid layer in the presence of uniform internal heat source. The effect of non-uniform basic temperature gradients on the onset of ferroconvection has been analyzed (Shivakumara et al. [14], and Shivakumara and Nanjundappa [15] [16]). Singh and Bajaj [17] have studied thermal convection of ferrofluids with boundary temperatures modulated sinusoidally about some reference value. Nanjundappa et al. [18] have studied the effect of internal heat generation on the criterion for the onset of convection in a horizontal ferrofluid saturated porous layer Nanjundappa et al. [19] have explored a model for penetrative ferroconvection via internal heat generation in a ferrofluid saturated porous layer. Nanjundappa et al. [20] have investigated the onset of penetrative Bénard-Marangoni convection in a horizontal ferromagnetic fluid layer in the presence of a uniform vertical magnetic field via an internal heating model. Ram and Kumar [21] has carried out to examine the effects of temperature dependent variable viscosity on the three dimensional steady axi-symmetric Ferrohydrodynamic (FHD) boundary layer flow of an incompressible electrically non conducting magnetic fluid in the presence of a rotating disk. Ram and Kumar [22] have analyzed the analysis of three dimensional rotationally symmetric boundary layer flow of field dependent viscous ferrofluid saturating porous medium. Ram et al. [23] have been made to describe the effects of geothermal viscosity with viscous dissipation on the three dimensional time dependent boundary layer flow of magnetic nanofluids due to a stretchable rotating plate in the presence of a porous medium. Ram et al. [24] have investigated numerically on the convective heat transfer behaviour of time-dependent three-dimensional boundary layer flow of nano-suspension over a radially stretchable surface. Kumar et al. [25] have studied the Bodewadt flow of a magnetic nanofluid in the presence of geothermal viscosity. Very recently, Ram et al. [26] have studied the rheological effects due to oscillating field on time dependent boundary layer flow of magnetic nanofluid over a rotating disk.

The purpose of this paper is to study the penetrative Bénard-Marangoni convection in a micropolar ferromagnetic fluid layer via internal heat generation. Such a study helps in understanding control of convection due to a non-uniform temperature gradient arising due to an internal heat source, which is important in the applications of ferrofluid technology. The linear stability problem is solved numerically using the Galerkin method, and the results are presented graphically. Moreover, the stability of the system when heated from below and also in the absence of thermal buoyancy is discussed in detail.

2. Mathematical Formulation

We consider an initially quiescent horizontal incompressible micropolar ferrofluid layer of characteristic thickness d in the presence of an applied uniform magnetic field H_0 in the vertical direction with the angular momentum $\boldsymbol{\omega}$. Let $T_0(z=0)$ and $T_1 < T_0(z=d)$ be the temperatures of the lower and upper rigid boundaries, respectively with $\Delta T (= T_0 - T_1)$ being the temperature difference. A uniformly distributed overall internal heat source is present within the micropolar ferrofluid layer. The Cartesian co-ordinate system (x, y, z) is used with the origin at the bottom of the layer and z-axis is directed vertically upward. Gravity acts in the negative z-direction, $\boldsymbol{g} = -g\hat{k}$ where \hat{k} is the unit vector in the z-direction.

The upper free boundary is assumed to be flat and subjected to linearly temperature dependent surface tension σ is $\sigma = \sigma_0 - \sigma_T (T - T_0)$, σ_T is the rate of thermal surface tension.

The governing equations for the flow of an incompressible micropolar ferromagnetic fluid are:

 ∇

$$\cdot \boldsymbol{q} = 0 \tag{1}$$

(4)

$$\rho_0 \left[\frac{\partial \boldsymbol{q}}{\partial t} + (\boldsymbol{q} \cdot \nabla) \boldsymbol{q} \right] = -\nabla p + \rho \boldsymbol{g} + (\boldsymbol{B} \cdot \nabla) \boldsymbol{H} + (\eta + \xi_r) \nabla^2 \boldsymbol{q} + 2\xi_r (\nabla \times \boldsymbol{\omega})$$
(2)

$$\rho_0 I \left[\frac{\partial \boldsymbol{\omega}}{\partial t} + (\boldsymbol{q} \cdot \nabla) \boldsymbol{\omega} \right] = \mu_0 \left(\boldsymbol{M} \times \boldsymbol{H} \right) + \nabla \left(\nabla \cdot \boldsymbol{\omega} \right) + \eta' \left(\nabla^2 \boldsymbol{\omega} \right) + 2\xi_r \left[\left(\nabla \times \boldsymbol{q} \right) - 2\boldsymbol{\omega} \right]$$
(3)

$$k_{1}\nabla^{2}T + \delta(\nabla \times \boldsymbol{\omega}) \cdot \nabla T + \boldsymbol{Q}''$$

= $+\mu_{0}T \left(\frac{\partial \boldsymbol{M}}{\partial T}\right)_{V,H} \cdot \frac{D\boldsymbol{H}}{Dt} + \left[\rho_{0}C_{V,H} - \mu_{0}\boldsymbol{H} \cdot \left(\frac{\partial \boldsymbol{M}}{\partial T}\right)_{V,H}\right] \frac{DT}{Dt}$

$$\rho = \rho_0 \left[1 - \alpha \left(T - T_0 \right) \right] \tag{5}$$

$$\nabla \cdot \boldsymbol{B} = 0, \ \nabla \times \boldsymbol{H} = 0 \text{ or } \boldsymbol{H} = \nabla \phi$$
 (6)

$$\boldsymbol{B} = \mu_0 \left(\boldsymbol{M} + \boldsymbol{H} \right) \tag{7}$$

$$\boldsymbol{M} = \frac{\boldsymbol{H}}{H} \boldsymbol{M} \left(\boldsymbol{H}, \boldsymbol{T} \right) \tag{8}$$

$$M = M_0 + \chi \left(H - H_0 \right) - K \left(\overline{T} - T_0 \right)$$
(9)

The basic state is assumed to be quiescent and is given by

$$\left[\boldsymbol{q}_{b},\boldsymbol{\omega}_{b},\boldsymbol{\rho},T,\boldsymbol{H},\boldsymbol{M}\right] = \left[0,0,\boldsymbol{\rho}_{b}\left(z\right),T_{b}\left(z\right),\boldsymbol{H}_{b}\left(z\right),\boldsymbol{M}_{b}\left(z\right)\right]$$
(10)

Using Equation (10) in Equation (2) and (4) respectively yield

$$\frac{\mathrm{d}p_b}{\mathrm{d}z} = -\rho_0 \Big[1 - \alpha_t \left(T_b - T_0 \right) \Big] g\hat{k} + \mu_0 M_b \frac{\mathrm{d}H_b}{\mathrm{d}z} \tag{11}$$

$$\frac{\mathrm{d}^2 T_b}{\mathrm{d}z^2} = -\frac{Q}{k_1} \tag{12}$$

Solving Equation (12) subject to the boundary conditions $T_b=T_0$ at z=0and $T_b=T_0-\Delta T$ at z=d, we obtain

$$T_{b}(z) = -\frac{Qz^{2}}{2k_{1}} + \frac{Qdz}{2k_{1}} - \beta z + T_{0}$$
(13)

Substituting Equation (6) after using Equations (9) and (13), the basic state magnetic field intensity $H_b(z)$ and magnetization $M_b(z)$ are found to be (see Finlayson [4])

$$\boldsymbol{H}_{b}(z) = \left[\boldsymbol{H}_{0} - \frac{K}{1+\chi} \left(\frac{Qz^{2}}{2k_{1}} - \frac{Qdz}{2k_{1}} + \beta z\right)\right]\hat{k}$$
(14)

$$\boldsymbol{M}_{b}(z) = \left[\boldsymbol{M}_{0} + \frac{K}{1+\chi} \left(\frac{Qz^{2}}{2k_{1}} - \frac{Qdz}{2k_{1}} + \beta z\right)\right]\hat{k}$$
(15)

where $M_0 + H_0 = H_0^{ext}$.

Using Equations (13) and (14) in Equation (11) and integrating, we obtain

$$p_{b}(z) = p_{0} - \rho_{0}gz - \rho_{0}\alpha g \left[\frac{Qz^{3}}{6k_{1}} - \frac{Qdz^{2}}{4k_{1}} + \frac{\beta z^{2}}{2} \right] - \frac{\mu_{0}M_{0}K}{1+\alpha} \left[\frac{Qz^{2}}{2k_{1}} - \frac{Qdz}{2k_{1}} + \beta z \right] - \frac{\mu_{0}K^{2}}{(1+\alpha)^{2}} \left[\frac{Q^{2}z^{4}}{8k_{1}^{2}} + \frac{z^{3}}{2} \left(\frac{Q\beta}{k_{1}} - \frac{Q^{2}d}{2k_{1}^{2}} \right) + \frac{z^{2}}{2} \left(\beta^{2} + \frac{Q^{2}d^{2}}{4k_{1}^{2}} - \frac{Q\beta d}{k_{1}} \right) \right]$$
(16)

The pressure distribution is of no consequence here as we are eliminating the same. It may be noted that $T_b(z)$, $H_b(z)$ and $M_b(z)$ are distributed parabolically with the porous layer height due to the presence of internal heat generation. However, when Q=0 (*i.e.*, in the absence of internal heat generation), the basic state temperature distribution is linear in z. Thus the presence of internal heat generation plays a significant role on the stability of the system.

To study the stability of the system, we perturb all the variables in the form

$$\begin{bmatrix} \boldsymbol{q}, \boldsymbol{\omega}, \boldsymbol{\rho}, \boldsymbol{p}, \boldsymbol{T}, \boldsymbol{H}, \boldsymbol{M} \end{bmatrix}$$

= $\begin{bmatrix} \boldsymbol{q}', \boldsymbol{\omega}', \boldsymbol{\rho}_b(z) + \boldsymbol{\rho}', \boldsymbol{p}_b(z) + \boldsymbol{p}', \boldsymbol{T}_b(z) + \boldsymbol{T}', \boldsymbol{H}_b(z) + \boldsymbol{H}', \boldsymbol{M}_b + \boldsymbol{M}' \end{bmatrix}$ (17)

where $q', \omega', \rho', p', T', H'$ and M' are the perturbed quantities and are assumed to be very small. Substituting Equation (17) into Equation (6) and using Equations (8) and (9) and assuming $K\beta d \approx (1+\chi)H_0$ and

 $KQd^2 \approx 2\kappa (1+\chi)H_0$ as propounded by Finlayson [4], we obtain (after dropping primes)

$$H_{x} + M_{x} = (1 + M_{0}/H_{0})H_{x},$$

$$H_{y} + M_{y} = (1 + M_{0}/H_{0})H_{y},$$

$$H_{z} + M_{z} = (1 + \chi)H_{z} - KT$$
(18)

where, (H_x, H_y, H_z) and (M_x, M_y, M_z) are the (x, y, z) components of the magnetic field and magnetization respectively. Thus the analysis is restricted to physical situation in which the magnetization induced by the variations in temperature gradient and internal heating is small compared that induced by external magnetic field.

Substituting Equation (17) into Equation (2), linearizing, eliminating the pressure term by operating curl twice and using Equations (18) the z-component of the resulting equation can be obtained as (after dropping the primes)

$$\begin{bmatrix} \rho_0 \frac{\partial}{\partial t} - (\eta + \xi_r) \nabla^2 \end{bmatrix} \nabla^2 w$$

$$= 2\xi_r \nabla^2 \Omega_3 \rho_0 \alpha g \nabla_1^2 T \left[\mu_0 K \nabla_1^2 \left(\frac{\partial \phi}{\partial z} \right) - \frac{\mu_0 K^2}{1 + \chi} \nabla_1^2 T \right] \left[\frac{Qz}{k_1} - \frac{Qd}{2k_1} + \beta \right]$$
(19)

Substituting Equation (17) into Equation (3) we obtain (after dropping primes)

$$\rho_0 I\left(\frac{\partial \Omega_3}{\partial t}\right) = -2\xi_r \left[\nabla^2 w + 2\Omega_3\right] + \eta' \nabla^2 \Omega_3 \tag{20}$$

As before, substituting Equation (17) into Equation (4) and linearizing, we obtain (after dropping primes)

$$\begin{bmatrix} \rho_0 C_0 \frac{\partial}{\partial t} - k_1 \nabla^2 \end{bmatrix} T = \begin{bmatrix} \rho_0 C_0 - \frac{\mu_0 T_0 K^2}{1 + \chi} \end{bmatrix} \begin{bmatrix} \underline{Qz} \\ k_1 - \underline{Qd} \\ 2k_1 + \beta \end{bmatrix} w + \mu_0 T_0 K \frac{\partial}{\partial t} \left(\frac{\partial \phi}{\partial z} \right) - \begin{bmatrix} \underline{Qz} \\ k_1 - \frac{Qd}{2k_1} + \beta \end{bmatrix} \partial \Omega_3$$
(21)

where $\rho_0 C_0 = \rho_0 C_{V,H} + \mu_0 H_0 K$.

Finally Equation (6), after using Equation (17) and (18), yield (after dropping primes)

$$\left(1+\chi\right)\frac{\partial^2\phi}{\partial z^2} + \left(1+\frac{M_0}{H_0}\right)\nabla_h^2\phi - K\frac{\partial T}{\partial z} = 0$$
(22)

Since the principle of exchange of stability is valid, the normal mode expansion of the dependent variables takes the form

$$w, T, \phi, \Omega_3 \} = \{ W(z), \Theta(z), \Phi(z), \Omega_3(z) \} \exp[i(lx + my)]$$
(23)

On non-dimensionalizing the variables by setting

$$\begin{pmatrix} x^*, y^*, z^* \end{pmatrix} = \begin{pmatrix} \frac{x}{d}, \frac{y}{d}, \frac{z}{d} \end{pmatrix}, \quad W^* = \frac{d}{\nu}W,$$

$$\Theta^* = \frac{\kappa}{\beta \nu d} \Theta, \quad \Phi^* = \frac{(1+\chi)\kappa}{\kappa\beta\nu d^2} \Phi,$$

$$\Omega^*_3 = \frac{d^3}{\nu}\Omega_3, \quad I^* = \frac{1}{d^2}I.$$

$$(24)$$

Equation (23) is substituted into Equations (19)-(22) and then Equation (24) is used to obtain the stability equations in the following form

$$(1+N_1)(D^2-a^2)^2 W$$

$$= a^2 R_t \Theta - 2N_1 (D^2-a^2) \Omega_3 - a^2 R_m [1+N_s (2z-1)](D\Phi-\Theta)$$
(25)

$$2N_{1}\left[\left(D^{2}-a^{2}\right)W+2\Omega_{3}\right]-N_{3}\left(D^{2}-a^{2}\right)\Omega_{3}=0$$
(26)

$$(D^{2} - a^{2})\Theta + [N_{s}(2z - 1) + 1][(1 - M_{2})W - N_{5}\Omega_{3}] = 0$$
(27)

$$D^2 \Phi - a^2 M_3 \Phi - D\Theta = 0 \tag{28}$$

The typical value of M_2 for magnetic fluids with different carrier liquids turns out to be of the order of 10^{-6} and hence its effect is neglected when compared to unity.

The above equations are to be solved subject to the rigid-paramagnetic boundary conditions:

$$W = DW = \Omega_3 = \Theta = \Phi = 0 \quad \text{at } z = 0$$

$$W = D^2 W + a^2 M a \Theta = D\Omega_3 = 0 \quad \text{at } z = 1$$

$$D\Theta + Bi\Theta = D\Phi = 0 \quad \text{at } z = 1$$
(29)

3. Numerical Solution

Equations (25)-(28) together with boundary conditions (29) constitute an eigenvalue problem with thermal Rayleigh number R_t being an eigenvalue. Accordingly, W, Θ, Φ and Ω_3 are written as

$$W(z) = \sum_{i=1}^{N} A_{i}W_{i}(z), \ \Omega_{3} = \sum_{i=1}^{N} B_{i}\Omega_{3i}(z)$$

$$\Theta(z) = \sum_{i=1}^{N} C_{i}\Theta_{i}(z), \ \Phi(z) = \sum_{i=1}^{N} D_{i}\Phi_{i}(z)$$
(30)

where A_i, B_i, C_i and D_i are the unknown constants to be determined. The basis functions $W_i(z)$, $\Theta_i(z)$, $\Phi_i(z)$ and $\Omega_{3i}(z)$ are generally chosen such that they satisfy the corresponding boundary conditions but not the differential equations. Substituting Equation (30) into Equations (25)-(28) and multiplying the resulting momentum Equation (25) by $W_j(z)$, angular momentum Equation (26) by $\Omega_{3j}(z)$, energy Equation (27) by $\Theta_j(z)$ and magnetic potential Equation (28) by $\Phi_j(z)$, performing integration by parts with respect to z between z = 0 and z = 1 and using the boundary conditions (29) we obtain the following system of linear homogeneous algebraic equations:

$$C_{ji}A_i + D_{ji}B_i + E_{ji}C_i + F_{ji}D_i = 0$$
(31)

$$G_{ji}A_i + H_{ji}D_i = 0 \tag{32}$$

$$I_{ji}A_i + J_{ji}B_i + T_{ji}D_i = 0 ag{33}$$

$$K_{ji}B_i + L_{ji}C_i = 0 \tag{34}$$

where the co-efficient C_{ji} - L_{ji} involve the inner product of the basis functions and are given by

$$\begin{split} C_{ji} &= (1+N_1) \Big[\left\langle D^2 W_j D^2 W_i \right\rangle + a^4 \left\langle W_j W_i \right\rangle + 2a^2 \left\langle D W_j D W_i \right\rangle \Big] \\ D_{ji} &= -a^2 R_i M_1 \left\langle \big[N_s \left(2z - 1 \right) + 1 \big] W_j \Theta_i \right\rangle - a^2 R_i \left\langle W_j \Theta_i \right\rangle + a^2 Ma D W_j (1) \Theta_i (1) \\ &= E_{ji} = a^2 R_i M_1 \left\langle \big[N_s \left(2z - 1 \right) + 1 \big] W_j D \Phi_i \right\rangle \\ &= F_{ji} = -2 N_1 \Big[\left\langle D W_j D \Omega_{3i} \right\rangle + a^2 \left\langle W_j \Omega_{3i} \right\rangle \Big] \\ &= G_{ji} = 2 N_1 \Big[\left\langle D \Omega_{3j} D W_i \right\rangle + a^2 \left\langle \Omega_{3j} W_i \right\rangle \Big] \\ H_{ji} &= -\Big[4 N_1 \left\langle \Omega_{3j} \Omega_{3i} \right\rangle + N_3 \left\langle D \Omega_{3j} D \Omega_{3i} \right\rangle + N_3 a^2 \left\langle \Omega_{3j} \Omega_{3i} \right\rangle \Big] \\ &= I_{ji} = (1 - M_2) \left\langle \big[N_s \left(2z - 1 \right) + 1 \big] \Theta_j W_i \right\rangle \\ &= J_{ji} = - \Big[\left\langle D \Theta_j D \Theta_i \right\rangle + a^2 \left\langle \Theta_j \Theta_i \right\rangle + Bi/4 \Big] \\ &= T_{ji} = -N_5 \left\langle \big[N_s \left(2z - 1 \right) + 1 \big] \Theta_j \Omega_{3i} \right\rangle \\ &= K_{ji} = a \Big[\Phi_j (1) \Phi_i (1) + \Phi_j (0) \Phi_i (0) \Big] \\ &\quad + \left\langle D \Phi_j D \Phi_i \right\rangle + a^2 M_3 \left\langle \Phi_j \Phi_i \right\rangle \end{split}$$

where the inner product is defined as $< > \langle ... \rangle = \int_{0}^{1} (...) dz$. The set of homogeneous algebraic equations can have non-trivial solutions if and only if

$$\begin{vmatrix} C_{ji} & D_{ji} & E_{ji} & F_{ji} \\ G_{ji} & 0 & 0 & H_{ji} \\ I_{ji} & J_{ji} & 0 & T_{ji} \\ 0 & K_{ji} & L_{ji} & 0 \end{vmatrix} = 0$$
(35)

The eigenvalue has to be extracted from the above characteristic equation. In Galerkin method, we choose the weighting function as the trial functions, thus:

$$W_{i} = z^{2} (z-1)^{2} z^{i-1}, \ \Theta_{i} = z (1-z/2) z^{i-1}, \ \Omega_{3i} = z (1-z/2) z^{i-1}, \ \Phi_{i} = z (1-z/2) z^{i-1}$$
(36)

The velocity (W_i) , temperature (Θ_i) , vorticity (Ω_{3i}) and magnetic potential (Φ_i) trail functions satisfy all the boundary condition while the temperature (Θ_i) does not satisfy the boundary condition $D\Theta + Bi\Theta = 0$ at z = 1. Therefore, following, the boundary residual technique is used for these functions. The velocity, vorticity and the magnetic equations are made orthogonal to each of the corresponding trail functions. For the temperature trial the boundary residuals are added and their combined inner product is set to zero to obtain $\langle D\Theta_j D\Theta_i \rangle + a^2 \langle \Theta_j \Theta_i \rangle + Bi\Theta_j (1)\Theta_i (1)$. Besides, the residual from this condition is included as residual from the differential Equation (36) leads to a relation involving in the form

$$f(R_t, R_m, Ma, N_s, M_1, M_3, N_1, N_3, N_5, a) = 0.$$

The critical values of R_t (*i.e.*, R_c) or R_m (*i.e.*, R_{mc}) or Ma (*i.e.*, Ma_c) is determined numerically with respect to *a* for different values of N_s , M_1 , M_3 , N_1 , N_3 and N_5 .

4. Result and Discussion

The classical linear stability analysis has been carried out to investigate the effect of internal heat source strength on the onset of Bénard-Marangoni ferroconvection in a horizontal micropolar ferrofluid layer heated from below in the presence of a transverse uniform vertical magnetic field. The both the boundaries is considered to be rigid-ferromagnetic. The critical thermal Rayleigh number (R_{tc}) , critical magnetic Rayleigh number (R_{mc}) and critical Marangoni number (Ma_c) and the corresponding critical wave number (a_c) are used to characterize the stability of the system. The critical stability parameters computed numerically by Galerkin technique as explained above, are found to converge by considering nine terms in the Galerkin expansion.

To validate the solution computed numerically for various values of R_t and Biin the absence of micropolar effects and internal heat source strength (*i.e.* $N_1 = N_3 = N_5 = Ns = 0$) are compared in **Table 1** with the previously published results of Davis [27]. In addition, the present method are compared with the previously published results of Char and Chiang [28] when $N_1 = N_3 = N_5 = 0$ and $R_m = R_t M_1 = 0$ (classical Rayleigh-Bénard problem) for various values of Ns (see **Table 2**). From the Tables, it is observed that our results are identical

	В	i = 0		<i>Bi</i> = 10		
	Davis [27]	Present Study		Davis [27]	Present Study	
R_{t}	Ma _c	Ma _c	R_{t}	Ma _c	Ma _c	
0.0	79.61	79.59	0.0	413.4	413.29	
100.0	68.43	68.47	100.0	378.7	378.62	
200.0	57.12	57.10	300.0	305.0	304.91	
300.0	45.49	45.48	500.0	225.1	225.08	
400.0	33.59	33.58	700.0	138.6	138.62	
500.0	21.39	21.38	900.0	44.73	44.729	

Table 1. Comparison of Ma_c for diff values of R_t and Bi in the absence of micropolar ferrofluid.

Table 2. Comparison of R_{tc} for diff values of *Ns* and *Bi* in the absence of micropolar ferrofluid.

Ns –	Char and Chiang [28]		Present study		Char and Chiang [28]		Present study	
	Free, isothermal		Free, isothermal		Free, insulated		Free, insulated	
	$(Bi \rightarrow \infty)$		$(Bi \rightarrow \infty)$		(<i>Bi</i> = 0)		(Bi = 0)	
	R_{tc}	a_{c}	R_{tc}	a_{c}	R_{tc}	a _c	R_{tc}	a_{c}
0	1100.684	2.682	1100.671	2.682	669.013	2.086	669.003	2.085
0.5	1055.612	2.679	1055.574	2.679	608.758	2.070	608.764	2.070
1	1011.471	2.680	1011.434	2.679	557.618	2.060	557.640	2.059
5	725.639	2.773	725.897	2.732	328.590	2.035	328.678	2.034
10	518.346	2.803	518.346	2.802	215.415	2.033	215.508	2.032
15	398.695	2.849	399.216	2.847	159.957	2.034	160.040	2.033
20	323.111	2.880	323.617	2.877	127.144	2.036	127.217	2.034
30	233.637	2.916	234.077	2.914	90.116	2.038	90.174	2.036
40	182.746	2.937	183.125	2.935	69.778	2.039	69.825	2.038
70	110.369	2.967	110.627	2.965	41.598	2.042	41.628	2.040
100	79.020	2.980	79.2141	2.978	29.629	2.043	29.651	2.041

with those obtained by Davis [27] as well as Char and Chiang [28] using different approaches.

The presence of internal heating makes the basic temperature, magnetic field and magnetization distributions to deviate from linear to parabolic with respect to micropolar ferrofluid layer height which in turn have significant influence on the stability of the system. To assess the impact of internal heat source strength *Ns* on the criterion for the onset of ferroconvection, the distributions of dimensionless basic temperature, $T_b(z)$, magnetic field intensity, $H_b(z)$ and magnetization, $M_b(z)$ are exhibited graphically in Figure 1 for various values of *Ns*. From the figure it is observed that increase in *Ns* amounts to large deviations in these distributions which in turn enhance the disturbances in the horizontal



Figure 1. Basic state temperature, magnetic intensity and magnetization distributions for different *Ns.*

porous layer and thus reinforce instability on the system.

Figures 2-4 depict the critical Ma_c at the onset of ferroconvection as the function of "*a*". It is noted that, as "*a*" decreases the Marangoni number decreases, attains a minimum at some critical wave number, and increases again. The curves reported in figures have the shape is upward concave to that of Bénard-Marangoni-ferroconvection. For increasing R_m , N_s , N_3 , R_ρ and decreasing N_1 is shifted to the neutral curves are slanted towards the higher wave number region.

Figure 5 represents the variation of critical Marangoni number Ma_c as a function of N_1 for different values of R_m and N_5 for $N_3 = 2$, $M_3 = 5$ and $N_s = 2$. It is seen that Ma_c decreases with an increase in R_m and hence its effect is to hasten the onset of ferroconvection due to an increase in the destabilizing magnetic force and the curve for $R_m = 0$ corresponds to non-magnetic micropolar fluid case. In other words, heat is transported more efficiently in magnetic fluids as compared to ordinary micropolar fluids. Also observed that Ma_c increases with increasing N_1 . This is because, as N_1 increases the concentration of microelements also increases and as a result a greater part of the energy of the system is consumed by these elements in developing gravitational velocities in the fluid which ultimately leads to delay in the onset of ferromagnetic convection. Moreover, the system is found to be more stable if the micropolar heat conduction of the parameter with $N_5 = 0.5$ as compared to the case of $N_5 = 0$.

In **Figure 6** Ma_c is plotted as a function of N_1 for different values of spin diffusion (couple stress) parameter N_3 and R_m when $M_3 = 5$, $N_5 = 0.5$ and Ns = 2. Here, it is observed that Ma_c curves for different N_3 coalesce when $N_1 = 0$. The impact of N_3 on the stability characteristics of the system is noticeable clearly with increasing N_1 and then it is seen that the critical Marangoni



Figure 2. Neutral curves for different values of R_m and N_5 with $N_1 = 0.5, R_t = 50, M_3 = 5, Ns = 2$.



Figure 3. Neutral curves for different values of R_i and N_1 for $N_3 = 2, N_5 = 0.5, R_m = 50, M_3 = 5$ and Ns = 2.

number decreases with increasing N_3 indicating the spin diffusion (couple stress) parameter N_3 has a destabilizing effect on the system. This may be attributed to the fact that as N_3 increases, the couple stress of the fluid increases, which leads to a decrease in micro-rotation and hence the system becomes more unstable.

Figure 7 shows the variation of critical Marangoni number Ma_c and as a function of N_1 for various values of dimensionless internal heat source strength Ns when $M_3 = 5$, $N_3 = 2$ and $N_5 = 0.5$. Figure 7 clearly indicates that Ma_c decreases monotonically with Ns indicating the influence of increasing internal heating is to decrease the value of Ma_c and thus destabilize the system. This is because increasing Ns amounts to increase in energy supply to the system.



Figure 4. Neutral curves for different values of N_3 and N_5 with $N_1 = 0.2, N_5 = 0.5, R_m = 50, R_t = 50$ and $M_3 = 5$.



Figure 5. Variation of Ma_c verses N_1 for different R_m fo $Ns = 2, M_3 = 5, N_3 = 2$.

The complementary effects of both buoyancy and magnetic forces are made clear in **Figure 8** by displaying the locus of Ma_c and magnetic Rayleigh number R_{mc} for various values of Bi and N_5 when $N_1 = 0.2$, $N_3 = 2$ and Ns = 2. We note that Ma_c is inversely proportional to R_{mc} due to the destabilizing magnetic force. From the figure it is evident that, increasing in Bi is to increase Ma_c and R_{mc} and thus its effect is to delay the onset of magnetic Bénard-Marangoni ferroconvection. This may be attributes to fact that with increasing Bi, the thermal disturbances can be easily dissipate in to the ambient surrounding due to a better convective heat transfer co-efficient at the top surface and hence higher



Figure 6. Variation of Ma_c verses N_1 for different N_3 for $N_5 = 0.5, Ns = 2, M_3 = 5$ and $R_t = 0$.



Figure 7. Variation of Ma verses N_1 for different Ns for $N_5 = 0.5, N_3 = 2, M_3 = 5$ and Bi = 0.

heating is required at make the system unstable. It is also evident that micropolar ferrofluid saturated porous layer in the presence of vertical magnetic field becomes more stable with increasing in *Bi*.

The measure of non-linearity of fluid magnetization M_3 , on the onset of ferroconvection is depicted in **Figure 9**. The curves of Ma_c versus R_{mc} shown in **Figure 9** for various values of M_3 when $N_3 = 2$, $N_5 = 0.3$, Ns = 2 and Bi = 2demonstrate that increasing M_3 has a destabilizing effect on the system. Nevertheless, the destabilization due to increase in M_3 is only marginal. This may be



Figure 8. Locus of Ma_c verses R_{mc} for different Bi and N_5 for $N_1 = 0.2, Ns = 2, M_3 = 1, N_3 = 2$ and $R_t = 50$.



attributed to the fact that the application of magnetic field makes the ferrofluid to acquire larger magnetization which in turn interacts with the imposed magnetic field and releases more energy to drive the flow faster. Hence, the system becomes unstable with a smaller temperature gradient as the value of M_3 increases. Alternatively, a higher value of M_3 would arise either due to a larger pyromagnetic coefficient or larger temperature gradient. Both these factors are conducive for generating a larger gradient in the Kelvin body force field, possibly promoting the instability.

5. Conclusions

The effect of internal heating and heat transfer coefficient on the onset of Bénard-Maranagoni-convection in a micropolar ferrofluid layer has been made theoretically. The solution of this problem is obtained numerically using Galer-kin-type of weighted residual technique by developing computer codes for MATTHEMAICA-11 software. Tabular and graphical method of appearance of the computed results illustrates the details in this paper and their dependence on the physical parameters involved in the problem. The significant findings of this analysis are:

1) The system becomes more unstable with an increase in magnetic Rayleigh number R_{nn} , nonlinearity of fluid magnetization parameter M_3 , internal heat source strength Ns and spin diffusion (couple stress) parameter N_3 .

2) The effect of increasing the value of coupling parameter N_1 , micropolar heat conduction parameter N_5 , Biot number *Bi* and is to delay the onset of ferromagnetic convection.

3) The effect of increasing R_m and Ns as well as decrease in N_1 , M_3 , N_3 and N_5 is to increase the critical wave number a_c and hence there is to reduce the convection cells.

4) The magnetic and buoyancy forces are complementary with each other and the system is more stabilizing when the magnetic forces alone are present.

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Nomenclature

q = (u, v, w): Velocity of the fluid *p*: Pressure *I* : Moment of inertia k_1 : Thermal conductivity *T* : Temperature C_{VH} : Specific heat at constant volume and magnetic field **B** : Magnetic induction field H : Magnetic field H_0 : Constant magnetic field $K = -(\partial M/\partial T)_{H_0,T_0}$: Pyromagnetic co-efficient **M** : Magnetization $M_0 = M(H_0, T_0)$: Constant mean value of magnetization Q'': Overall uniformly distributed effective volumetric internal heat generation H = |H| $M = |\mathbf{M}|$ D = d/dz: Differential operator $a = \sqrt{\ell^2 + m^2}$: Overall horizontal wave number $R_{i} = \alpha \beta g d^{4} / \nu \kappa$: Gravity thermal Rayleigh number $M_1 = \mu_0 K^2 \beta / (1 + \chi) \rho_0 \alpha g$: Magnetic number $M_2 = \mu_0 T_0 K^2 / (1 + \chi) \rho_0 C_0$: Magnetic parameter $M_3 = (1 + M_0/H_0)/(1 + \chi)$: Non-linearity of magnetization $R_m = R_t M_1$: Magnetic Rayleigh number $N_1 = \xi_r / \eta$: Coupling parameter: $N_3 = \eta' / \eta d^2$: Spin diffusion parameter $N_5 = \delta / \rho_0 C_0 d^2$: Micropolar heat conduction parameter $N_s = Qd/2k_1\beta$: Dimensionless heat source strength $P_r = \eta / \kappa$: Prandtl number **Greek Symbols** ρ : Density η : Shear kinematic viscosity co-efficient ξ_r : Vortex (rotational) viscosity $\boldsymbol{\omega}$: Angular velocity of colloidal particles along z-axis ρ_0 : Reference density μ_0 : Free space magnetic permeability

 η' : Shear spin viscosity co-efficient

 α : Thermal expansion co-efficient

 δ : Micropolar heat conduction coefficient

 $\chi = (\partial M / \partial H)_{H_0, T_0}$: Magnetic susceptibility

 ϕ : Magnetic potential

 $\nabla^2 = \partial^2 / \partial x^2 + \partial^2 / \partial y^2 + \partial^2 / \partial z^2$: Laplacian operator

 $D/Dt = \partial/\partial t + \mathbf{q} \cdot \nabla$: Convective derivative

 $\beta = \Delta T/d$ ($\Delta T > 0$): Temperature gradient